

Lecture 16: Gravitational potential energy and space travel

- Universal gravitational potential energy
- Space travel problems
- Escape speed
- Orbital energy
- Multiple objects

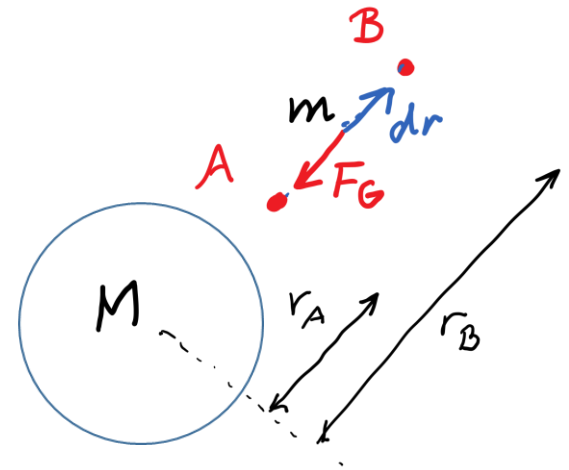
Gravitational potential energy

$$F_{grav} = \frac{GmM}{r^2}, \text{ attractive} \quad \text{Conservative force}$$

$$U_B - U_A = -W_{A \rightarrow B} = - \int_{\vec{r}_A}^{\vec{r}_B} \vec{F} \cdot d\vec{r}$$

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$$\int_{\vec{r}_A}^{\vec{r}_B} \vec{F} \cdot d\vec{r} = \int_{r_A}^{r_B} F_r dr = \int_{r_A}^{r_B} -\frac{GmM}{r^2} dr$$



$$= - \left[-\frac{GmM}{r} \right]_{r_A}^{r_B} = \frac{GmM}{r_B} - \frac{GmM}{r_A}$$

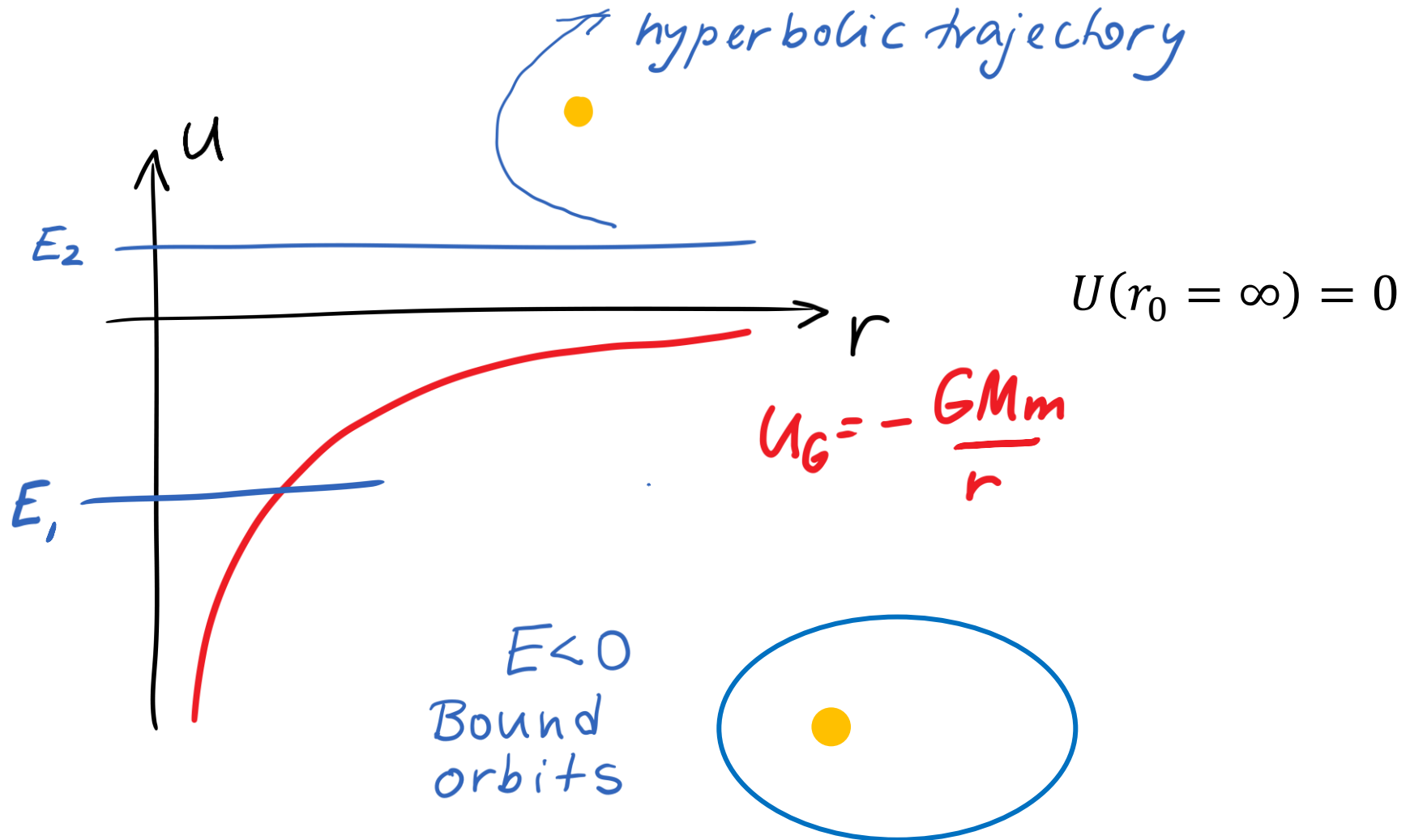
$$U_B - U_A = -GmM \left(\frac{1}{r_B} - \frac{1}{r_A} \right)$$

Choose reference point: $r_0 = \infty$.
Assign $U(r_0 = \infty) = 0$

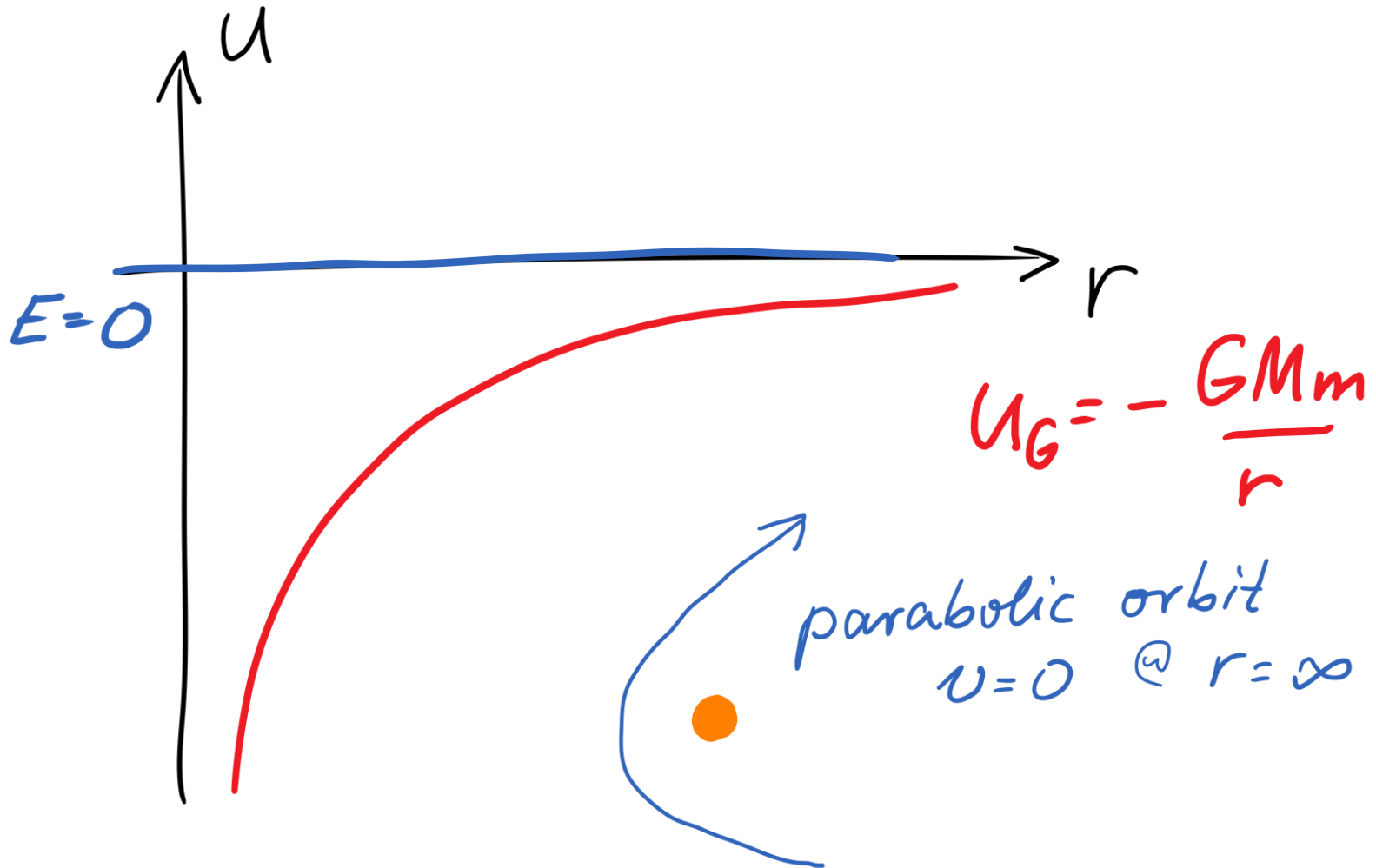
$$U_{\text{Grav}} = -\frac{GmM}{r}$$

negative!

Potential energy diagram

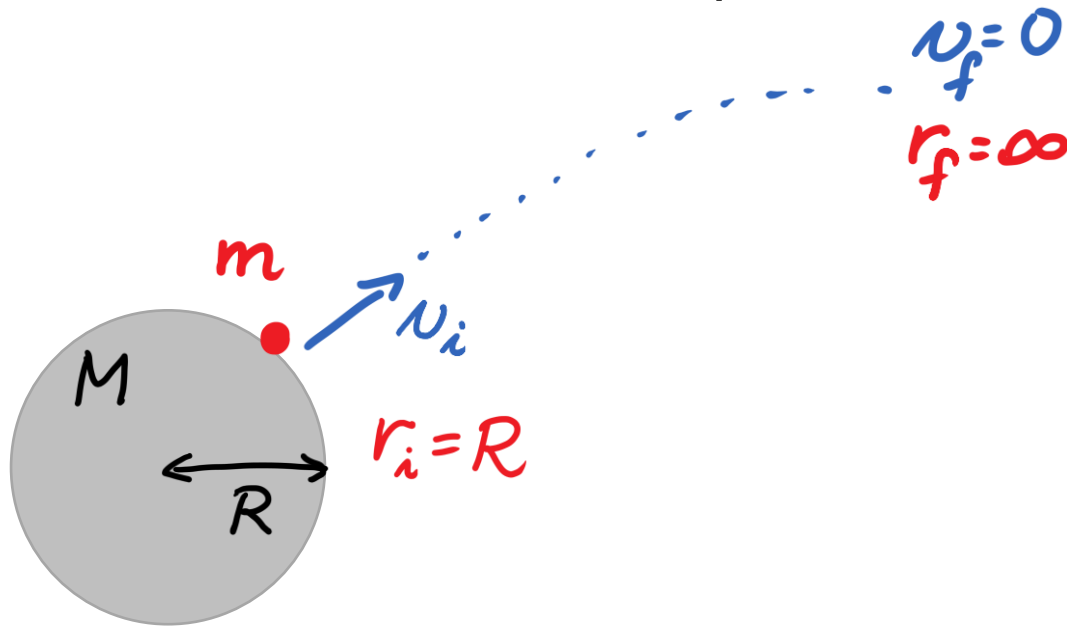


Potential energy diagram



Escape condition

$$E = \frac{1}{2}mv^2 - \frac{GmM}{r}$$

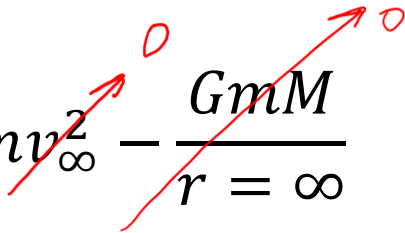


Critical escape condition: *barely* making it to $r = \textit{infinity}$, having slowed to speed of zero

Escape speed

Minimum speed an object must have at distance R from central mass M if it is to go infinitely far away

$$E_i = E_f$$

$$\frac{1}{2}mv_{esc}^2 - \frac{GmM}{R} = \cancel{\frac{1}{2}mv_{\infty}^2} - \cancel{\frac{GmM}{r = \infty}}$$


$$\cancel{\frac{1}{2}mv_{esc}^2} - \cancel{\frac{GmM}{R}} = 0 - 0$$

$$v_{esc} = \sqrt{\frac{2GM}{R}}$$

Example: Escape speed from Earth

$$M_{Earth} = 5.97 \times 10^{24} kg$$

$$R_{Earth} = 6.38 \times 10^6 m$$

$$G = 6.67 \times 10^{-11} N m^2/kg^2$$

$$v_{esc} = \sqrt{\frac{2GM}{R}} = 11,200 m/s$$

Example: Escape speed from orbit

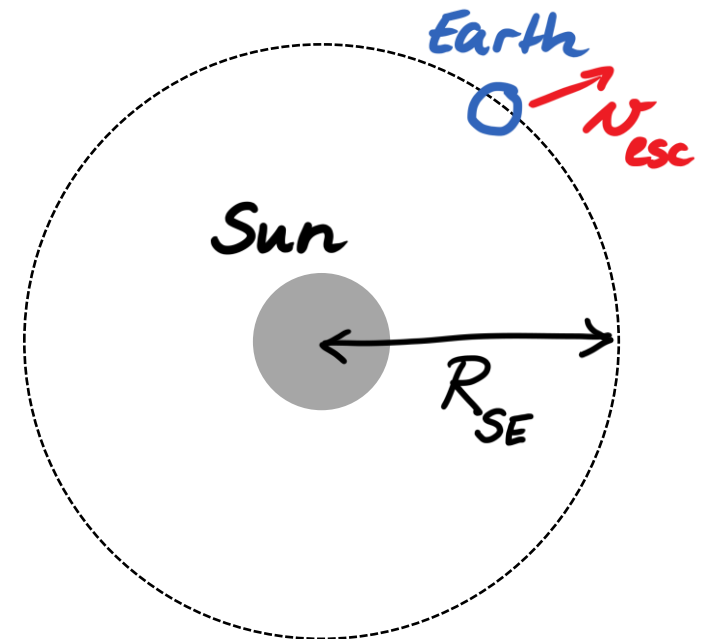
Speed necessary to escape gravitational field of the **Sun** when object is launched from Earth:

$$M_{Sun} = 1.99 \times 10^{30} \text{ kg}$$

$$R_{SE} = 1.50 \times 10^{11} \text{ m}$$

$$G = 6.67 \times 10^{-11} \text{ N m}^2/\text{kg}^2$$

$$v_{esc} = \sqrt{\frac{2GM_{Sun}}{R_{SE}}} = 42,000 \text{ m/s}$$



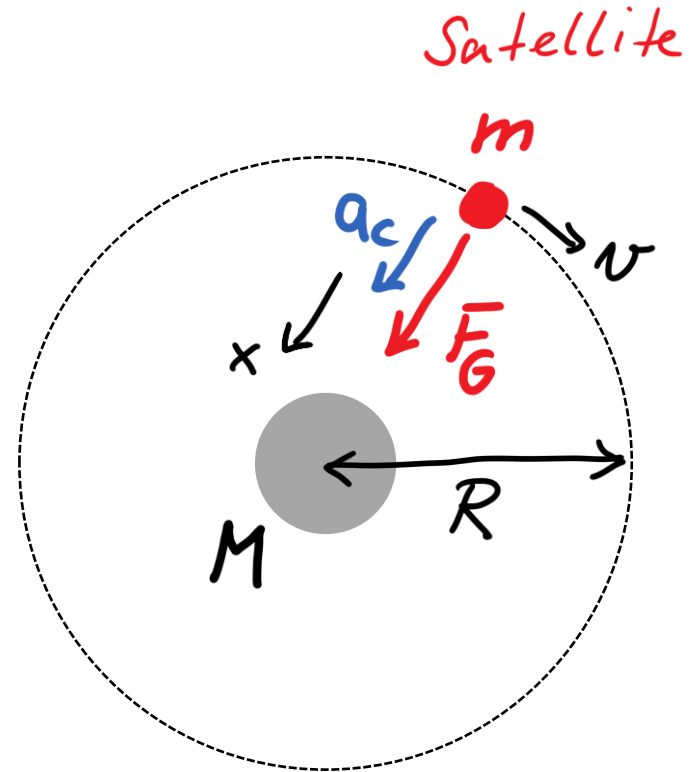
Orbital Energy

$$E = K + U = \frac{1}{2}mv^2 - \frac{GmM}{R}$$

Speed of satellite in
circular orbit: $v^2 = \frac{GM}{R}$

$$E = \frac{1}{2}m \left(\frac{GM}{R} \right) - \frac{GmM}{R}$$

$$E = -\frac{GmM}{2R}$$

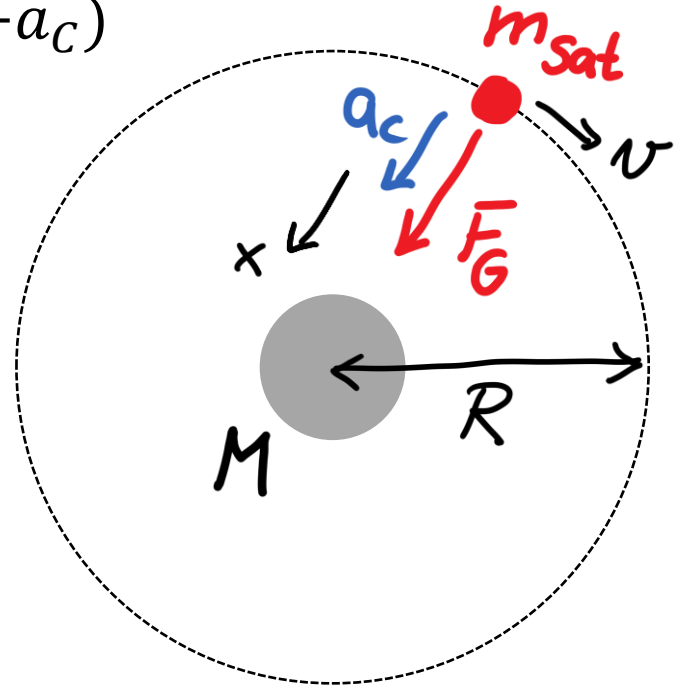


Satellite Motion

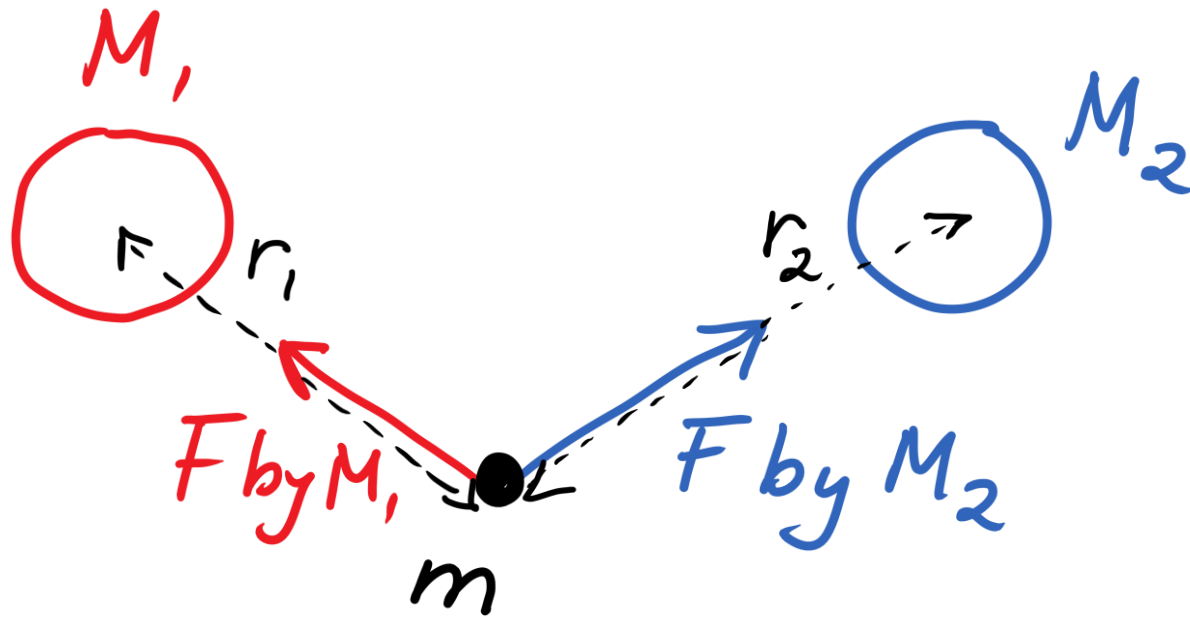
$$\sum F_x = F_{grav,x} = m_{sat} a_x = m_{sat} (+a_c)$$

$$\frac{Gm_{sat}M}{R^2} = m_{sat} \frac{v^2}{R}$$

$$\frac{GM}{R} = v^2$$



Multiple objects



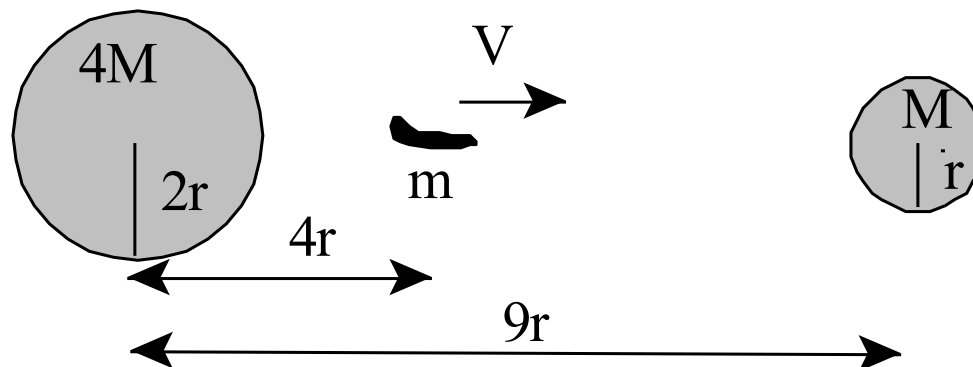
$$U_g = U_{g_1} + U_{g_2} = -\frac{GM_1 m}{r_1} + \left(-\frac{GM_2 m}{r_2}\right)$$

Example with multiple objects

A planet has mass $4M$ and a radius $2r$. Its moon has mass M and radius r . The centers of the planet and moon are a distance $9r$ apart.

A shuttle of mass m is a distance $4r$ away from the center of the planet and moving with speed V .

What is the total mechanical energy of the shuttle?



If the shuttle was initially at rest at X, how much work did the engines do?

